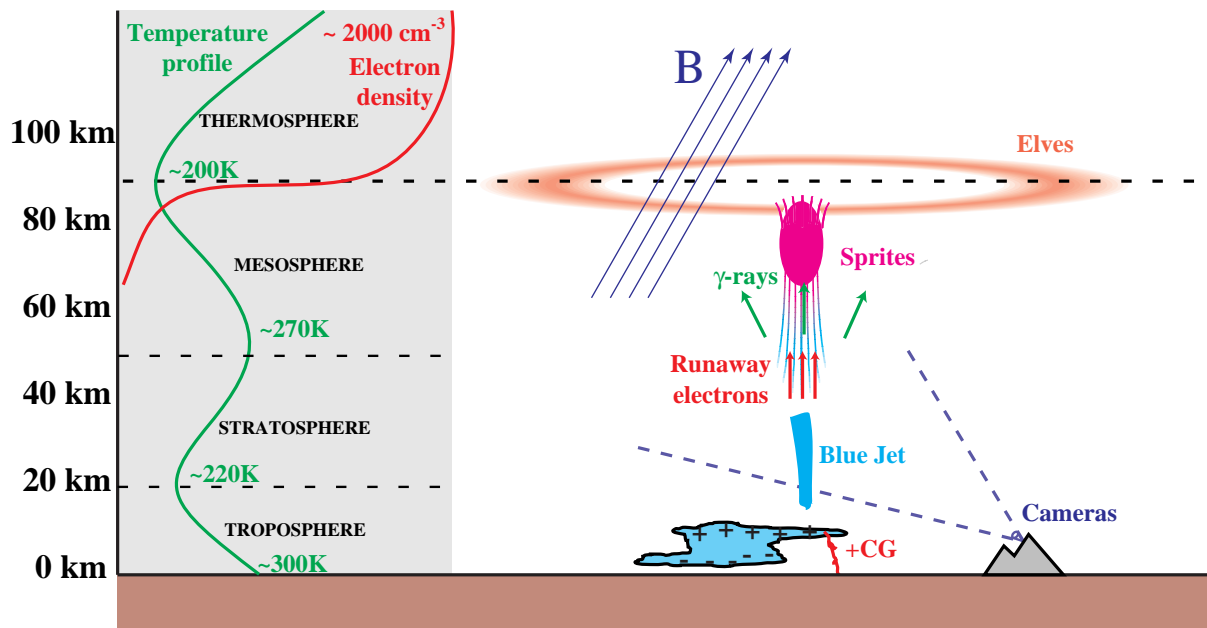


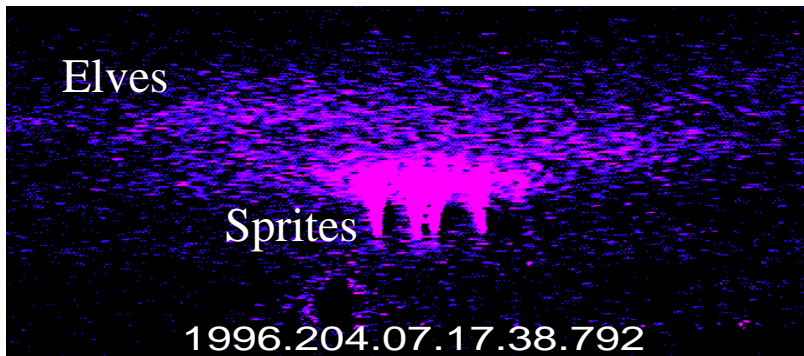
Monte Carlo Model of Runaway Electrons in the Mesosphere

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(STAR Laboratory, Stanford, CA 94305)*

Lightning-mesosphere interaction phenomena



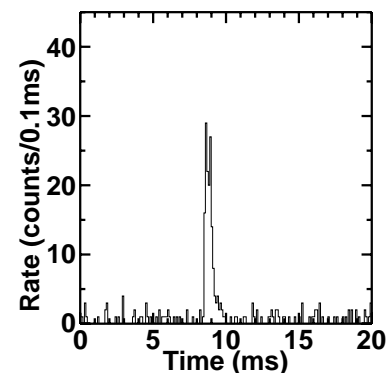
Red Sprites



Red Sprites:

- altitude range ~50-90 km
- lateral extent ~5-10 km
- occur ~1-5 ms after +CG discharge
- last up to several 10 ms

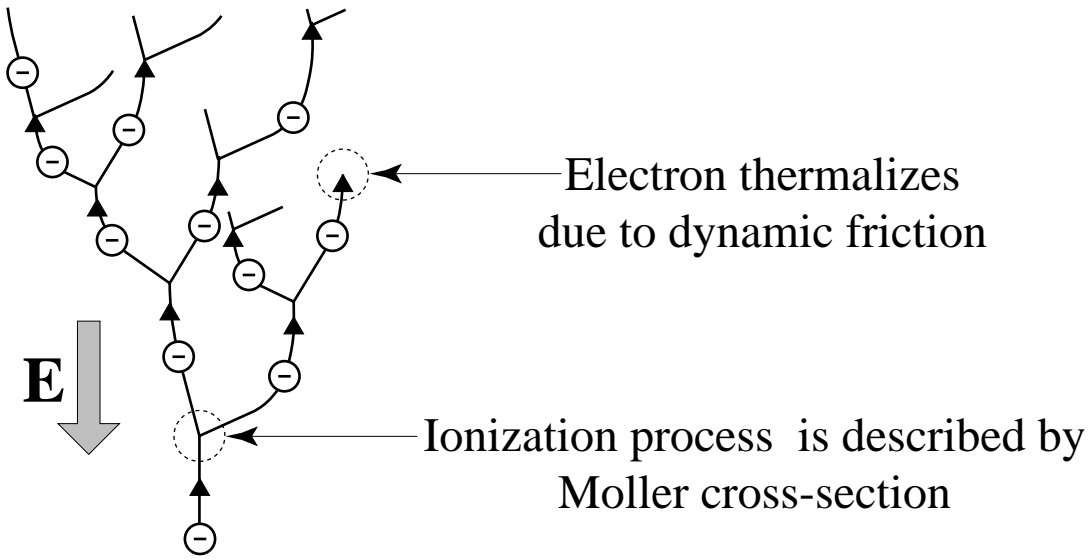
γ-ray flash
(BATSE observation)



Terrestrial γ-ray flashes:

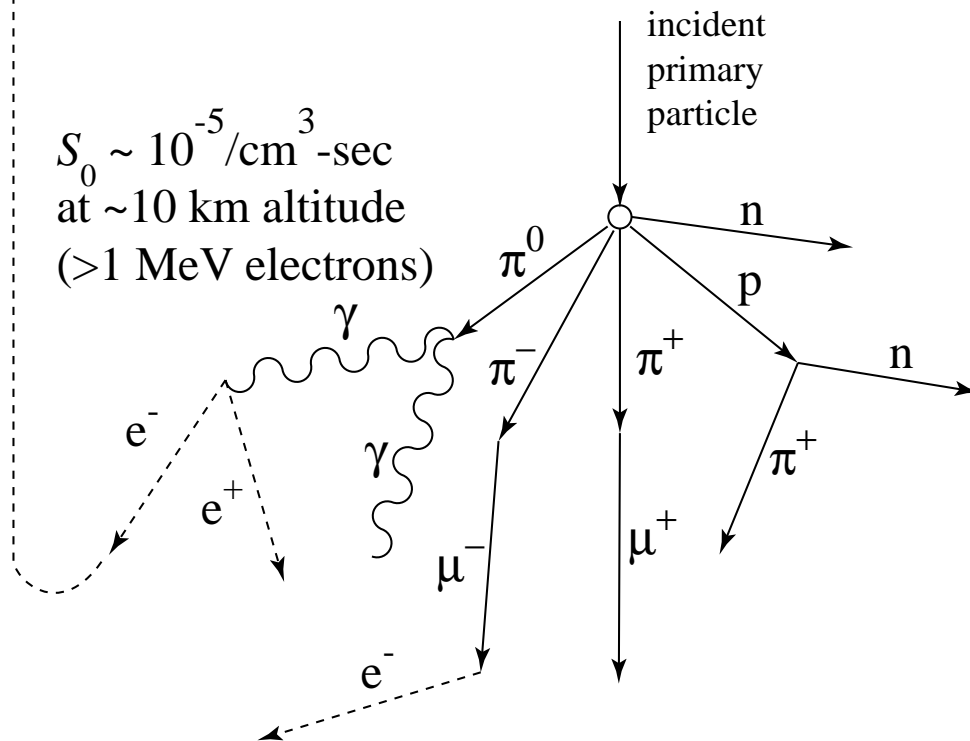
- last ~1 ms
 - were correlated with thunderstorms
- [Inan et al., 1996]

Runaway Electron Avalanche



Cosmic ray shower

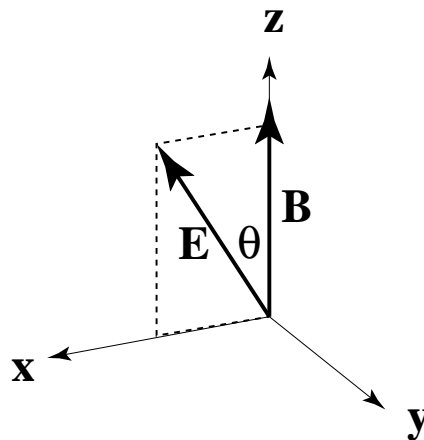
$S_0 \sim 10^{-5} / \text{cm}^3 \text{-sec}$
at $\sim 10 \text{ km}$ altitude
($>1 \text{ MeV}$ electrons)



Monte Carlo simulation

Previous runaway avalanche models:

- analytical [*Gurevich et al.*, 1996; *Sizykh et al.*, 1993; *Bulanov et al.*, 1997];
- kinetic [*Symbalisty et al.*, 1997];
- Monte Carlo [*Shveigert*, 1988].



Equation of motion:

$$\frac{d\vec{p}}{dt} = -e\vec{E} - \frac{e}{m\gamma}[\vec{p} \times \vec{B}] + \vec{\Gamma}(t)$$

Production of new electrons:

$$P(\vec{p} \text{ creates } \vec{p}' \text{ in interval } d^3\vec{p}') = N_m Z_m d\sigma dt$$

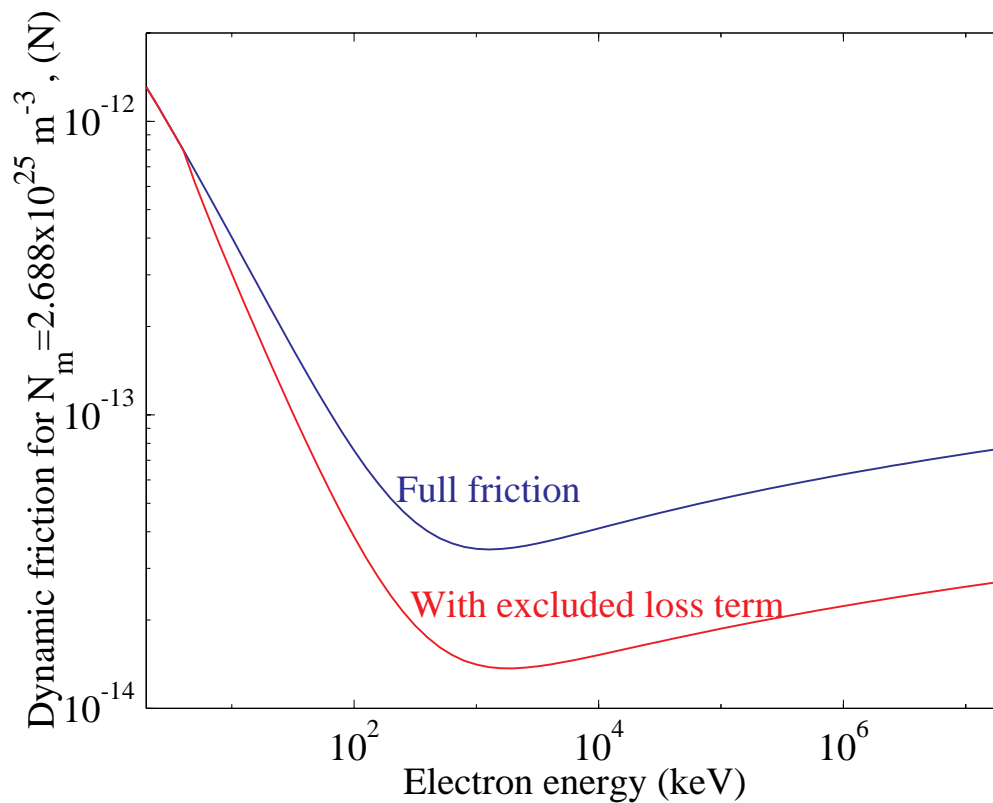
Angular diffusion part of $\Gamma(t)$:

$$\text{change direction by } \delta\theta = \sqrt{4D\Delta t}$$

Dynamic friction function

Exclude losses to created electrons from the friction:

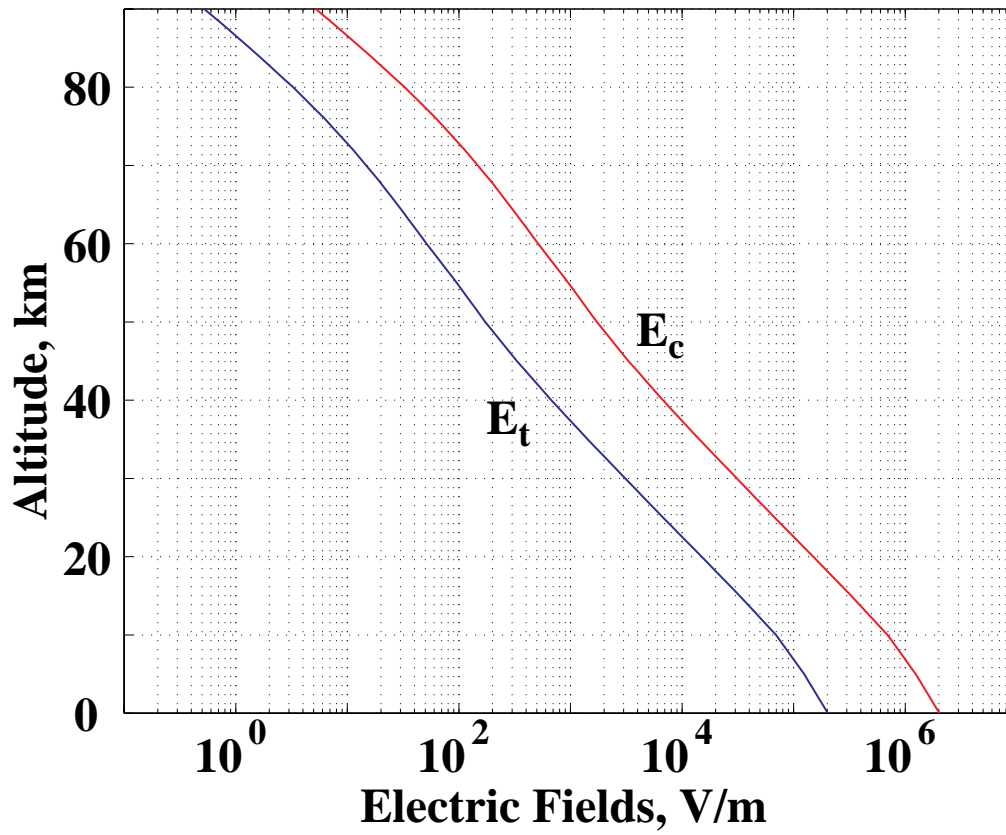
$$F_{d,\text{excl}}(p) = F_d(p) - N_m Z_m \int_{\mathcal{E}_{\min}}^{\mathcal{E} - \mathcal{E}_{\min}} \Delta \mathcal{E} \sigma(\mathcal{E}, \mathcal{E}') d\mathcal{E}'$$



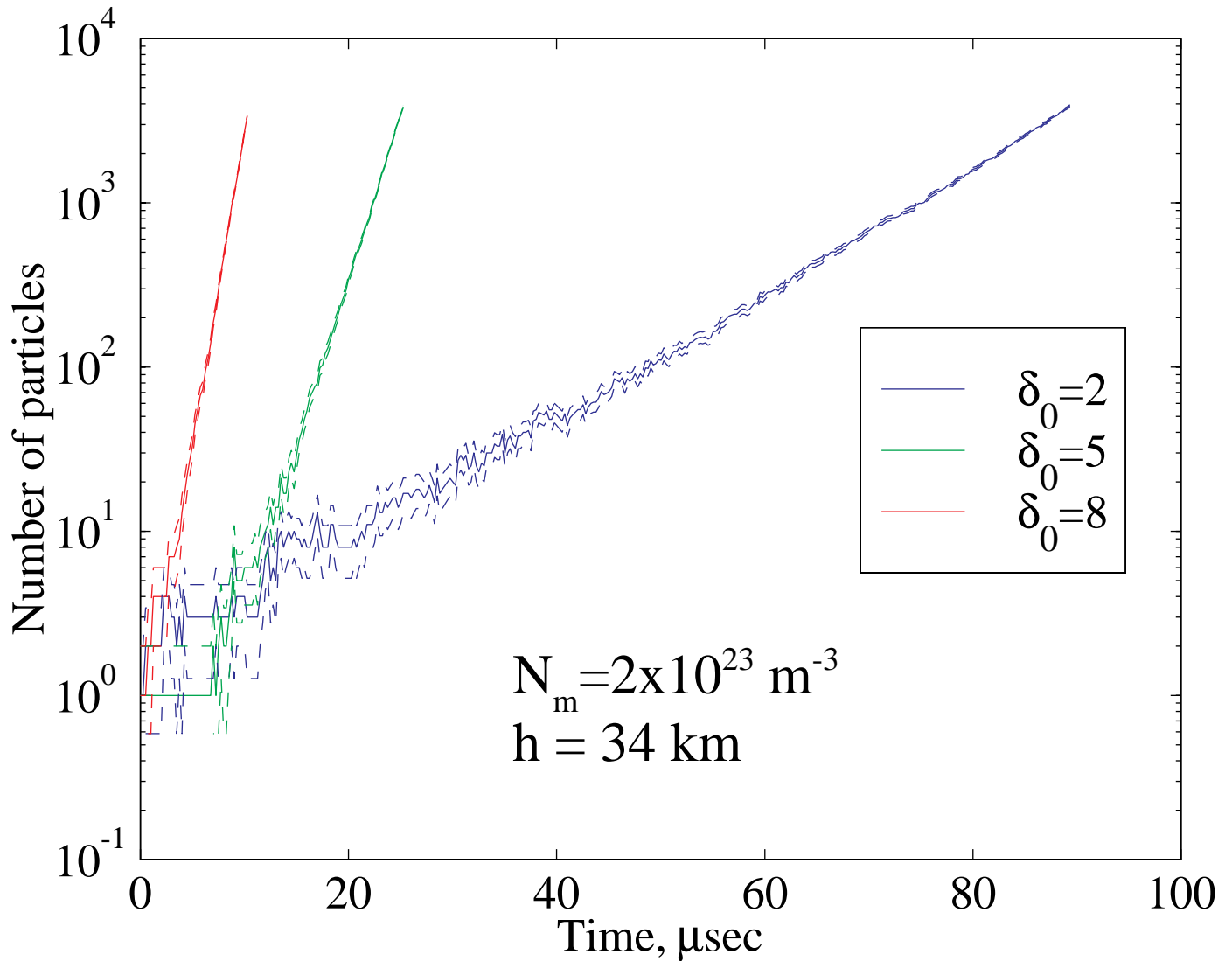
Runaway and conventional breakdown fields

$$F_d \sim N_m$$

$$E_t = F_{d,\min} / e$$



Growth of the number of particles for different parameters $\delta_0 = E/E_t$



Comparison of different predictions of runaway electron growth rate $1/\tau_i$

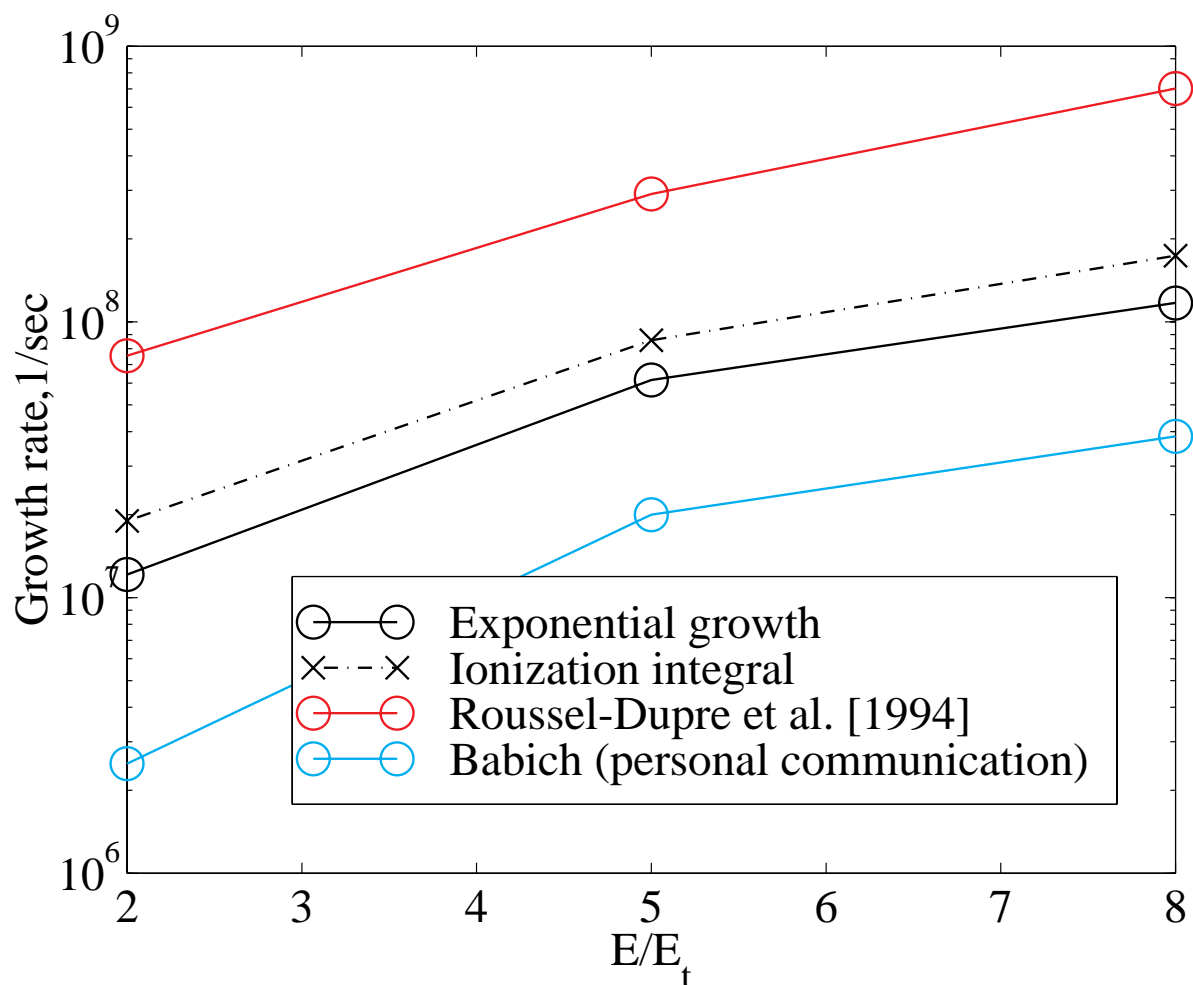
Two ways to calculate $1/\tau_i$:

1. Exponential growth of electron number:

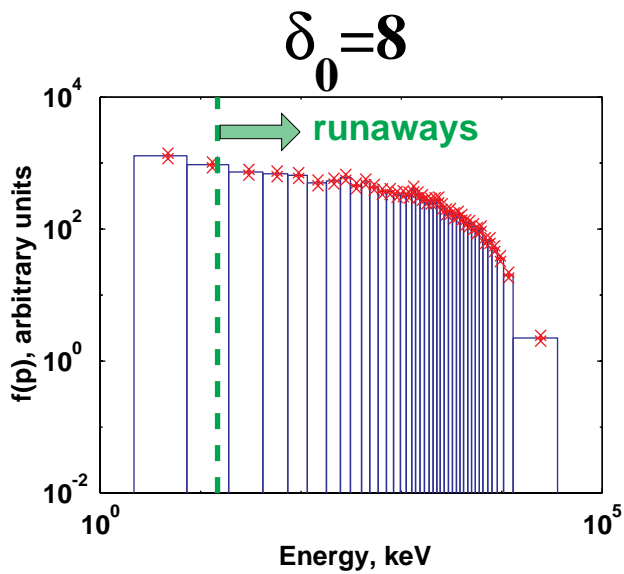
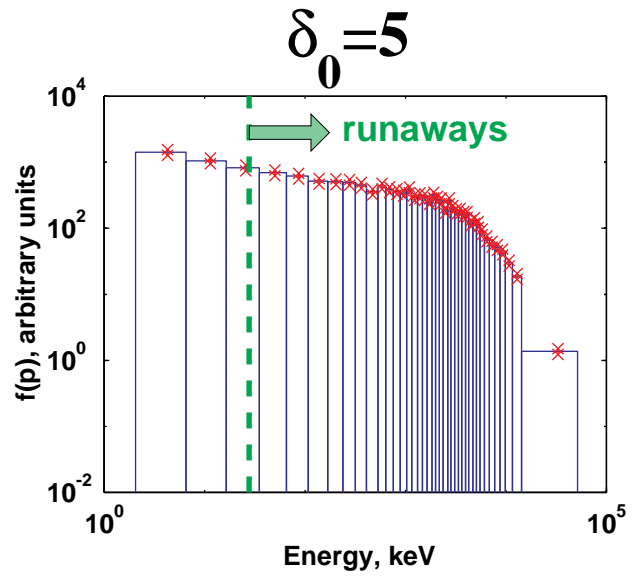
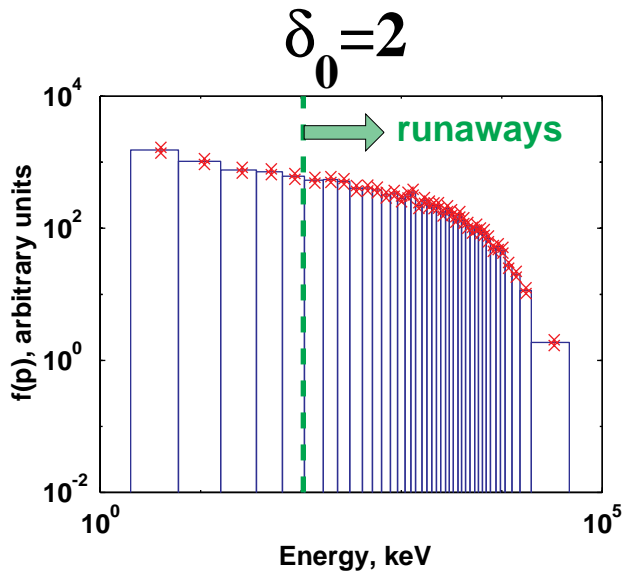
$$N_R(t) \sim e^{t/\tau_i}$$

2. Ionization integral:

$$\frac{1}{\tau_i} = \frac{\partial N_R / \partial t}{N_R} = \frac{\int_{\text{REL producers}} N_m v \sigma_{\text{tot}} f(\vec{p}) d^3 \vec{p}}{\int_{\text{REL}} f(\vec{p}) d^3 \vec{p}}$$

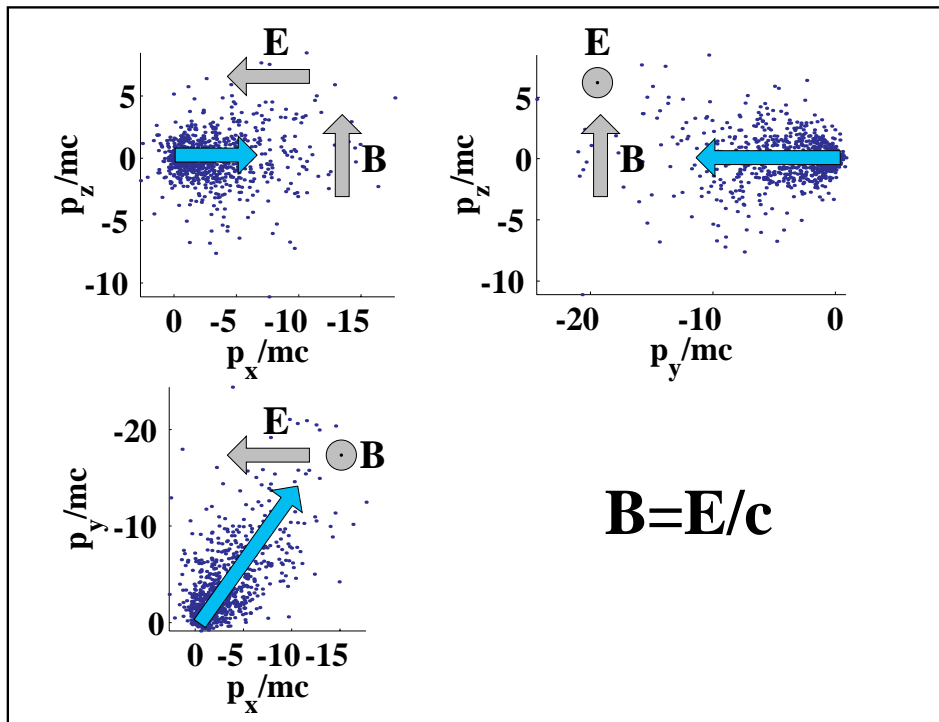
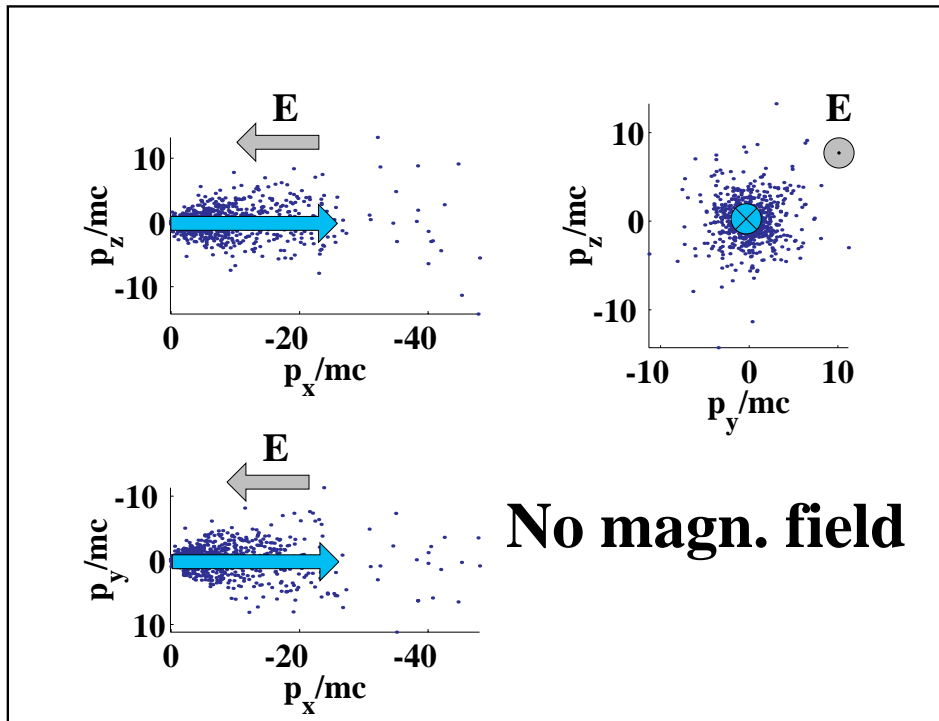


Electron distribution function without magnetic field for different $\delta_0 = E/E_t$



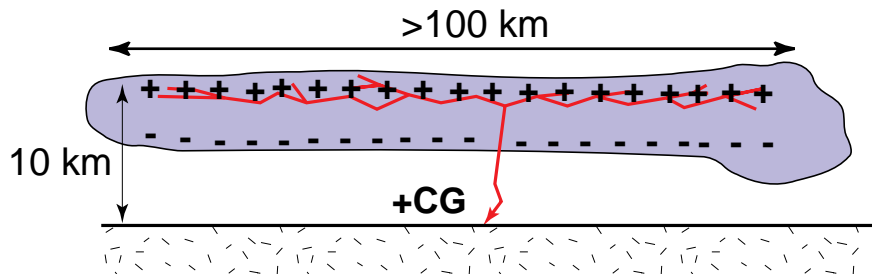
Electrons in momentum space for orthogonal E and B

$E=5E_t$



$B=2E/c$: No avalanche

Upper atmosphere model

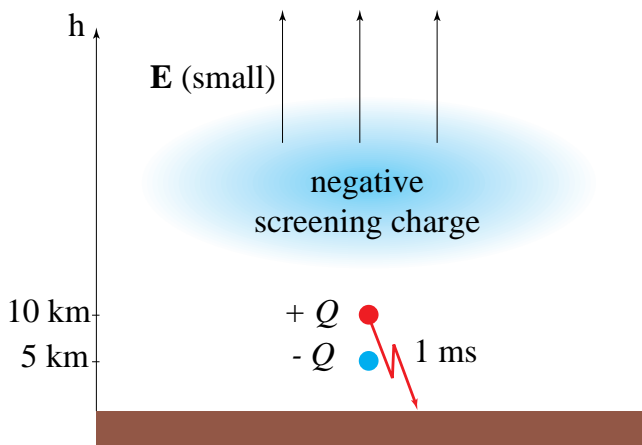


- cartesian (translationally symmetric along y axis)
- quasielectrostatic [*Pasko et al.*, 1997]
- fluid model for runaway electrons [*Lehtinen et al.*, 1997]

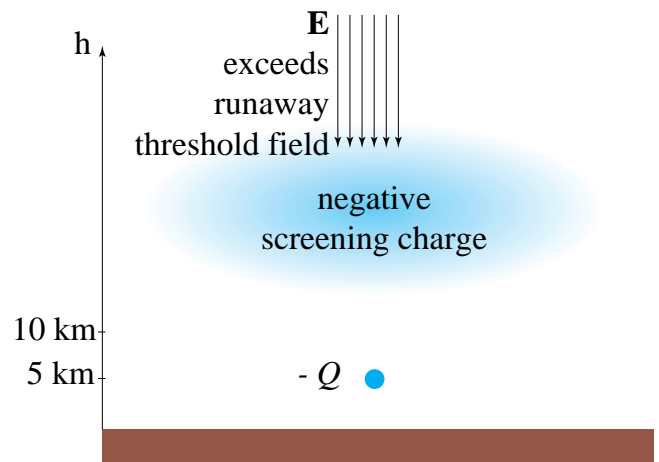
Dimensionless parameters for runaway velocity and rate:

$$\delta_0 = E/E_t, \quad \eta_0 = cB/E_t, \quad \mu_0 = \cos(\vec{E}, \vec{B})$$

BEFORE DISCHARGE

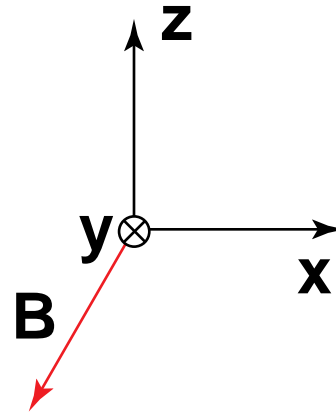
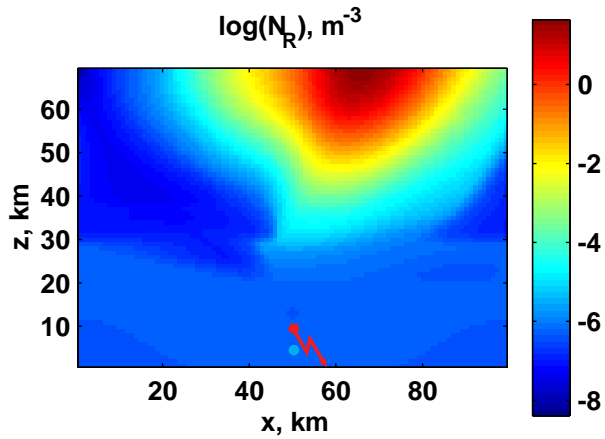
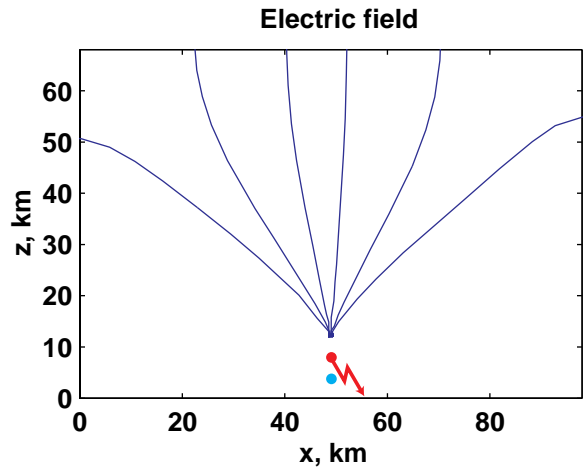
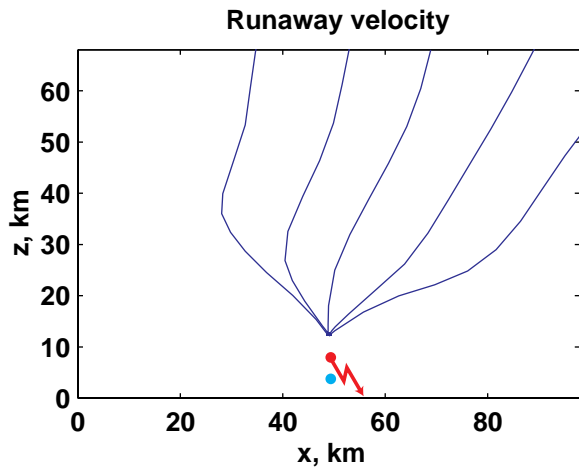


AFTER DISCHARGE



Runaway electron motion in presence of E and B: 1. B lies in (x,z) plane

$$Q = 8 \text{ C/km}$$

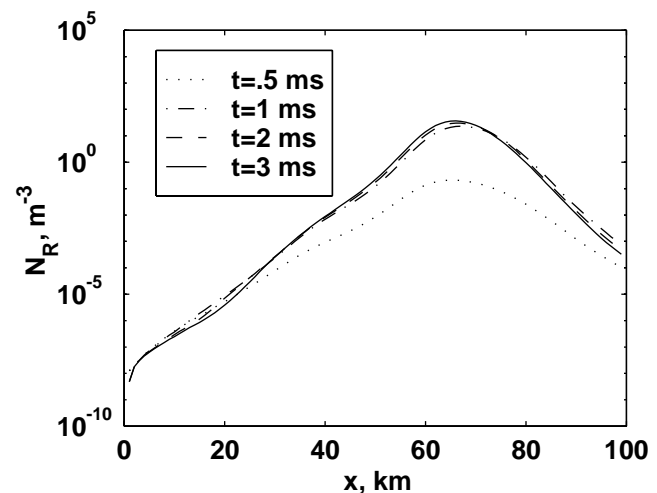
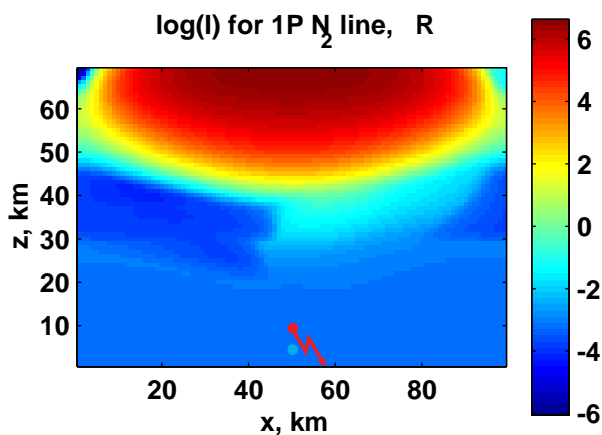
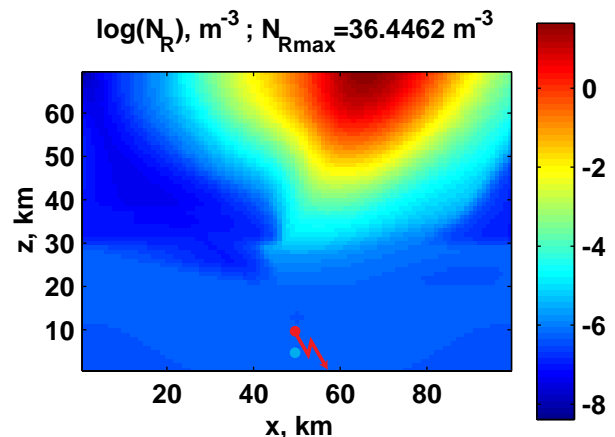
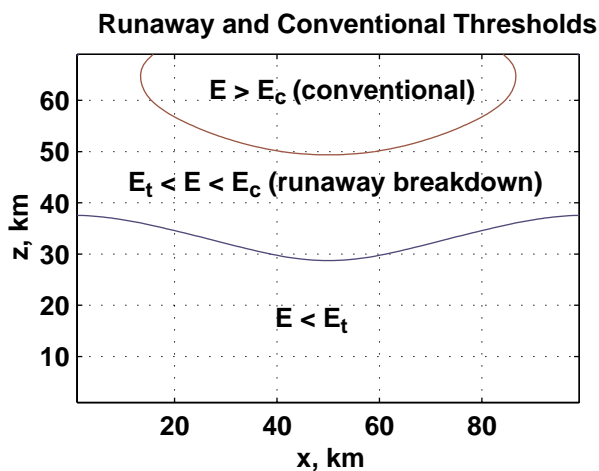
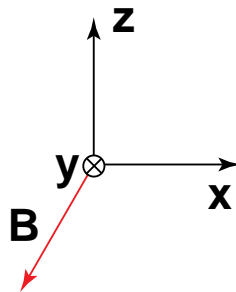


2D structure of runaway electron density and optical emissions in the First Positive Band of N_2 at $t = 3\text{ms}$:

Band of N_2 at $t = 3\text{ms}$:

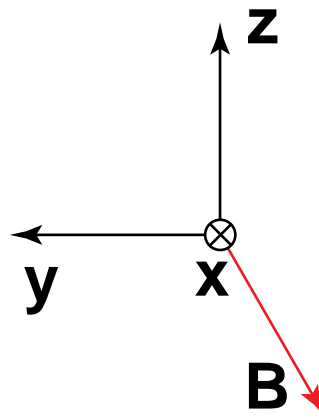
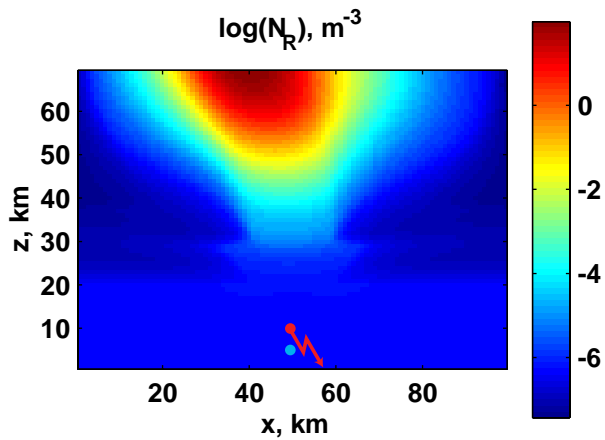
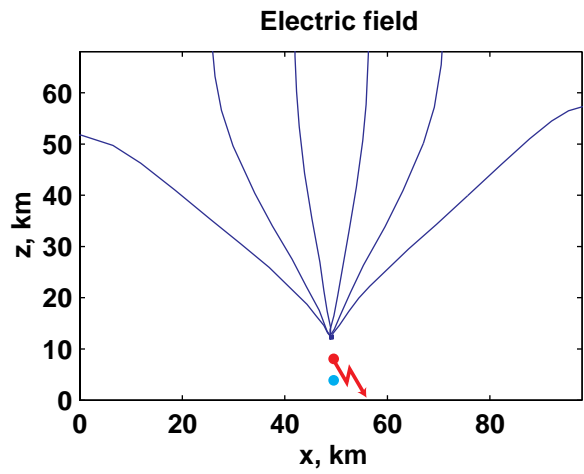
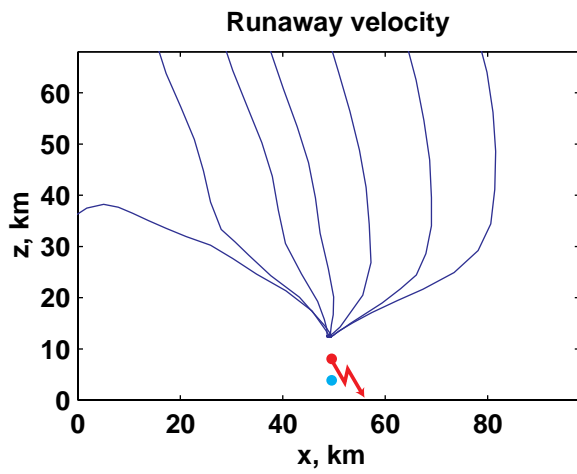
1. B lies in (x,z) plane

$$Q = 8 \text{ C/km}$$



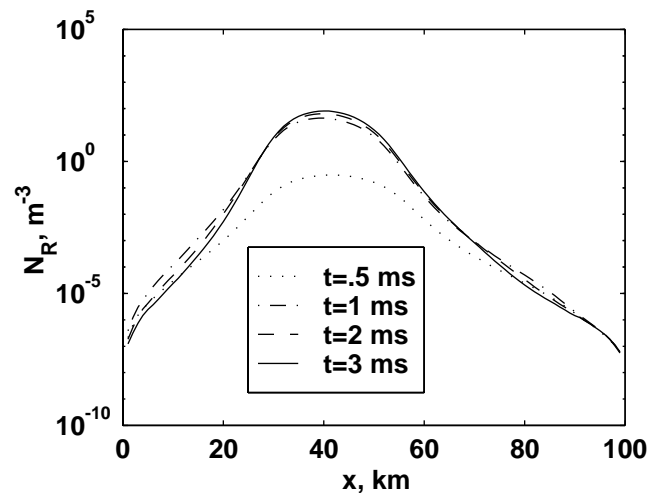
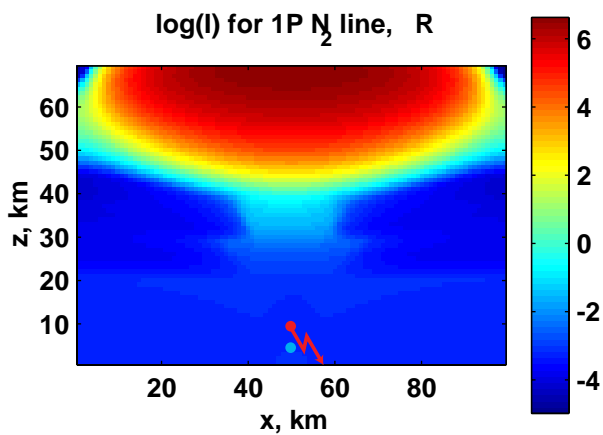
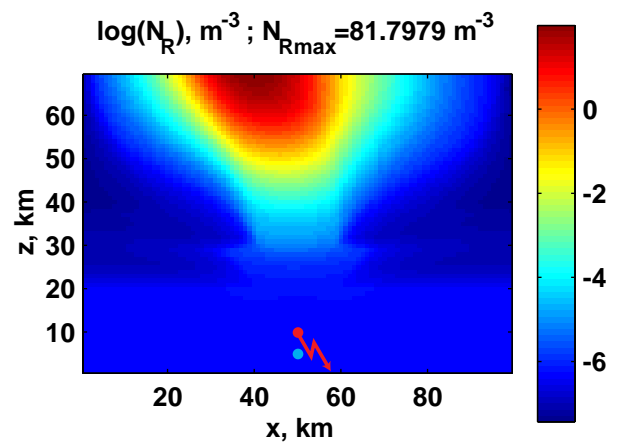
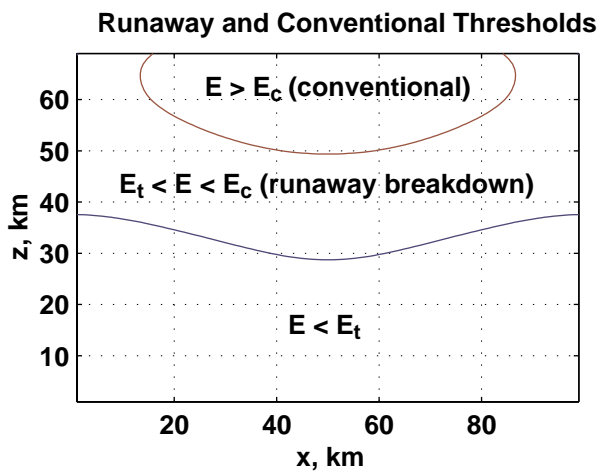
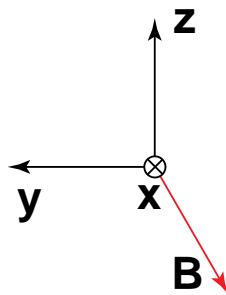
Runaway electron motion in presence of E and B: 2. B lies in (y,z) plane

$$Q = 8 \text{ C/km}$$

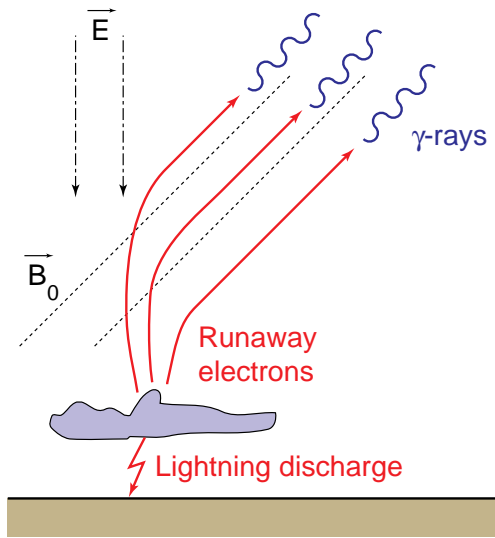


2D structure of runaway electron density and optical emissions in the First Positive Band of N_2 at $t = 3\text{ms}$: 2. B lies in (y,z) plane

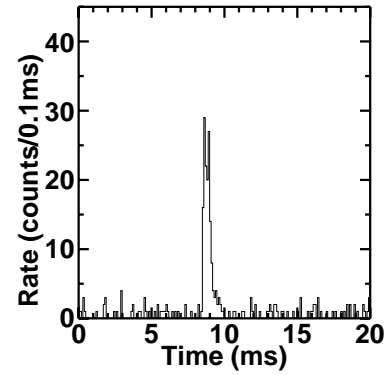
$$Q = 8 \text{ C/km}$$



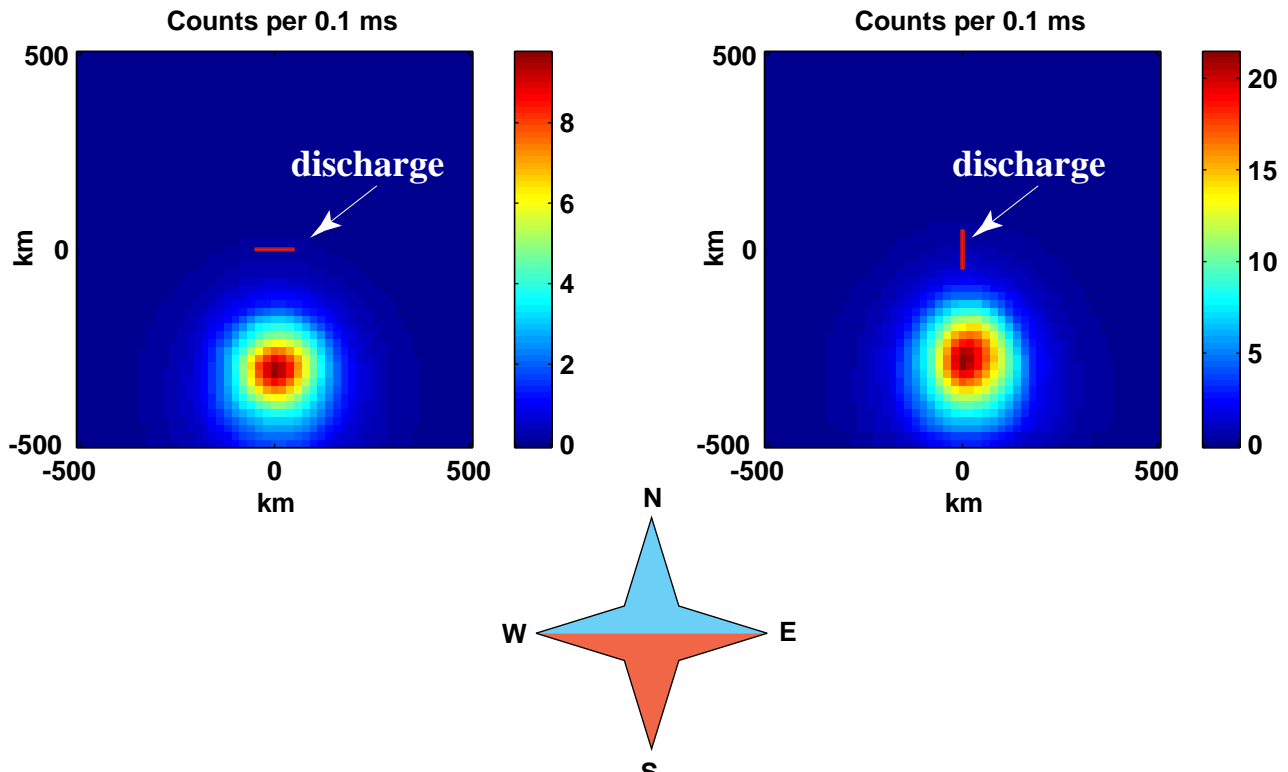
Terrestrial Gamma Ray Flashes



γ -ray flash
(BATSE observation)



Simulated BATSE data at $\sim 45^\circ$ magnetic N latitude in energy interval 100-300 keV



Conclusions

- We calculated uniform runaway electron avalanche rates in constant electric and magnetic fields and compared them to previously done work.
- We modelled runaway breakdown due to a positive return stroke from a laterally extensive thundercloud using cartesian (translationally symmetric) model and a lookup table of calculated runaway electron velocities and avalanche rates.
- The geomagnetic field controls the motion of runaways at $>35\text{km}$ at mid-latitudes, where most Sprites are observed, and close to equatorial region, where the terrestrial γ -ray flashes are observed.
- At mid-latitudes geomagnetic field doesn't retard the runaway electron avalanche since the angle between E and B is small.
- For sufficiently large discharge values, the REL-produced γ -rays flux values agree with BATSE data [*Fishman et al.*, 1994].

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